



# **MIST project @ MARI site**

### Ricardo Bustos Universidad Católica de la Santísima Concepción

2nd Global 21cm Workshop – McGill U., Montreal, 2019



# **MARI Project**

- MARI: Medidor Autónomo de Radio-Interferencia.
- Goal: Find a site with low RFI levels for redshifted 21cm observations.
- Measured RFI levels between 50 250 MHz in the north of Chile.
- Team:
  - Ricardo Bustos (PI, UCSC)
  - Raúl Monsalve (ASU at the time, now at McGill)
  - Judd Bowman (ASU)
  - Alexandra Suárez (UdeC)
  - Students (Electrical Engineering from UCSC)



UCSC









### MARI - STAGE I:





#### MARI Stage I results: vertical antenna, 1-hr integration





#### MARI Stage II - Characteristics

Deeper observations at the site with lowest RFI levels.

At a plateau we called the **MARI site**: west side of the Andes, blocking RFI from Chilean cities, at 4380 masl.



#### MARI Stage II results: July 2015, MARI site



### MARI site

#### Occupancy in [%] for 6677 spectra: 5 continuous days each month May, July, November



Bustos R., et al, 2019: In preparation



### MARI site – PWV studies 183 GHz radiometer - Oct. 2017

To evaluate the site for microwave frequencies, we measured PWV. We found 40% higher than the ALMA site, being 700 m lower. Still a competitive site at those frequencies



### MARI - June 2018

Weather station to characterize the MARI site (ESO funding)





# Legal status MARI site



Government permission:

# Antenna beam comparison Simulation v/s measurements

• Oscar Restrepo, PhD Student, Universidad de Chile.

#### EDGES Blade antenna scaled from 100 - 200 MHz to 1 - 2 GHz



Antenna simulation

Antenna panels: 0.6 mm

Antenna panels: 3.0 mm



# **Anechoic chamber measurements**

At UTFSM, Valparaíso.





# **Example from First Try**

More refined measurements in progress.



(b) Plano  $\phi = 90^{\circ}$ 



### MIST: Mapper of the IGM Spin Temperature

Goal:

- try to verify and measure with high precision global 21 cm signal.
- deployed and observing in 2020.

Team: Mauricio Díaz, Ricardo Bustos (UCSC), Raúl Monsalve, Jon Sievers, Cynthia Chiang (McGill), Nithyanandan Thyagarajan (NRAO), McGill, U. Chile, UCSC students.

Students from McGill: Christian Bye, Matheus Pessoa.



# Block Diagram



- Design based on EDGES-2, and somewhat similar to EDGES-3.
- All components different from EDGES
- Initial focus is on instrument, antenna not defined yet.
- Observation site not defined yet.



### Receiver



Temperature Control



### Instrument CAD Models









### Receiver Pictures





### **Receiver Pictures**





## Back End Pictures



Analog-to-Digital Converter



Switch Driver Board



Vector Network Analyzer



## Most of the Instrument





# Work Ahead

#### 1) Instrument and Calibration

- Finish assembly of the instrument.
- Finish the control program.
- Deep integration to test digitizer.
- Build receiver calibration unit.
- Lab receiver calibration at different temperatures.

#### 2) Antenna Design

• Simple antenna with smooth frequency response, sacrificing reflections and losses.

#### 3) Smaller is Better

- In parallel, continue development of portable version of MIST, "mini MIST".
- Small profile and low power, with enough performance for 21-cm science.



# http://mari-ucsc.blogspot.cl @Lab\_Astro\_UCSC

# Many thanks!

### MARI Stage I - Characteristics

- Antenna: Biconical dipole in horizontal (NS and EW) and vertical orientations.

- Frequency resolution: 36.3 kHz.
- Scan time: 3 min over range 50 250 MHz.
- Observing time per antenna orientation: 1-hour.

